$b \simeq c \simeq 5.7$ Å) (Landolt-Börnstein, 1969). Neither space group is a subgroup of the other, however, both are subgroups of $Pm3m-O_n^1$ which is also the space group of BaTiO₃ at high temperature. Thus the relevant group-subgroup relationships are $Pm3m \rightarrow R3m$ and $Pm3m \rightarrow Amm2$. Both, R3m and Amm2 are translationengleiche (though not maximal) subgroups of Pm3m and therefore only twin domains and no antiphase domains are allowed in the R3m and Amm2forms of BaTiO₃.

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Analytic Approximations for the Incoherent X-ray and Electron Intensities of Light Atoms and Ions

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It has been found that the incoherent scattered X-ray and electron intensities of light atoms, with 2 to 18 electrons, and of light ions, with 2 to 28 electrons, can be accurately approximated by an analytical formula containing six parameters.

The incoherent intensities of the first 36 atoms calculated by Tavard, Nicolas & Rouault (1967) have been fitted analytically by Hajdu (1971, 1972), those of the atoms of higher atomic numbers (from Ca to Am) were approximated by Pálinkás (1973) with the numerical values of Cromer (1967, 1969).

In these works, the validity of the approximations had been tested in the scattering-variable range available for X-ray diffraction, $0 \le s = \sin \theta / \lambda \le 1.5$ Å⁻¹.

In the present paper, a new analytical expression with six parameters is proposed for similar purposes. Its applicability is demonstrated in two somewhat different domains:

(1) for light atoms of atomic numbers 2 to 18 over a wide range of the scattering variable: $0 \le s \le 4.0$ Å⁻¹,

(2) for all the ions $(2 < Z \le 30)$ for which numerical incoherent-intensity tables were available. The scattering-variable range of these published tables extends to $s=1\cdot 1$ Å⁻¹, so the fit could not be checked beyond this limit in the case of the ions.

We endeavoured to find an approximating formula with good analytic behaviour. By the use of such a function, physical quantities determined by the incoherent scattering function, *e.g.* the exchange energy, and the contribution to the electron-electron correlation function (Bartell & Gavin, 1964), can also be expressed analytically.

We have found that an expression meeting these conditions and yielding a good approximation can be composed of terms similar to the function describing the incoherent scattering of the hydrogen atom:

$$I_{l}^{x}(S)/I_{T} \equiv S(s) = N \left[1 - \sum_{l=1}^{3} \frac{\gamma_{l}}{(1 + \lambda_{l}s^{2})^{n}} \right]$$
(1)

where $I_i^x(s)$ is the X-ray incoherent intensity, I_T is the Thomson formula, N is the number of electrons of the atom or the ion, γ_i and λ_i are independent parameters, and the power n of the denominator is 4 in the case of atoms and ions with $2 < Z \le 20$, and n = 1 for the ions with $21 \le Z \le 29$.*

The parameters γ_l and λ_l have been determined by a least-squares program, fitting the formula to the tabulated numerical intensities, *i.e.* the expression

$$\sum_{i=1}^{M} \delta_{i}^{2} = \sum_{i=1}^{M} [S_{fit}(s_{i}) - S_{tab}(s_{i})]^{2}$$

was minimized where *i* refers to the discrete $\sin \theta / \lambda$ points and *M* is the number of tabulated values.

The goodness of the fits is characterized in our tables by the percentage error defined as

$$\varepsilon = \frac{100}{S(s_M)} \sqrt{\frac{\sum\limits_{i=1}^{M} \delta_i^2}{M-1}}.$$
 (2)

Since the incoherent scattered electron intensities can be expressed – within the validity range of the first Born approximation – in terms of the corresponding X-ray intensity function (Morse, 1932), incoherent scattered electron intensities can also be expressed from (1) as

$$I_{t}^{e} = \frac{4}{a_{0}^{2}} \frac{S(s)}{(4\pi s)^{4}}$$
(3)

where a_0 is the Bohr radius.

Light neutral atoms with atomic number 2 to 18

Incoherent scattering functions of the atoms have been determined by Cromer (1967, 1969), based on the Waller-Hartree theory with the use of the numerical HF SCF wave functions of Mann (1967). On the basis of Cromer's tables we have obtained a good analytic fit for the atoms with 2 to 18 electrons by applying equation (1) with n=4. Table 1 contains the parameter values and the errors.

The accuracy of the calculations has been checked by a sum-rule procedure. It is known that the exchangeenergy term of the two-particle potential energy of the scatterer is related to the incoherent scattering function in the following way (Bartell & Gavin, 1964);

$$\frac{1}{\pi} \int_0^\infty [S(s) - N] \mathrm{d}s = \left\langle \frac{1}{r_{12}} \right\rangle_{\mathrm{exch}}.$$
 (4)

* For Zn^{2+} (Z=30) the better fit was obtained with n=4.

Table 1. Parameters $(\times 10^4)$ of the analytic fits of S(s)for atoms with electrons 2 to 18

| $n=4, s_{\max}=4.$ | | | | | | | | | |
|--------------------|----|------------|------|------------|-------------|-------------|-------------|------|--|
| | Ν | 7 1 | Y2 | 7 3 | λ_1 | λ_2 | λ_3 | з | |
| He | 2 | 3 | 7675 | 2321 | 940858 | 49201 | 24780 | 170 | |
| Li | 3 | 3506 | 901 | 5603 | 683877 | 15197 | 15196 | 2066 | |
| Be | 4 | 5180 | 0 | 4846 | 327829 | 104297 | 7693 | 2331 | |
| В | 5 | 4755 | 1304 | 3949 | 157768 | 157767 | 4762 | 902 | |
| С | 6 | 6437 | 198 | 3366 | 88239 | 375057 | 3258 | 289 | |
| Ν | 7 | 5756 | 1287 | 2958 | 47646 | 145022 | 2380 | 229 | |
| 0 | 8 | 6274 | 1106 | 2622 | 34741 | 120632 | 1799 | 323 | |
| F | 9 | 6532 | 1106 | 2362 | 25790 | 98417 | 1407 | 377 | |
| Ne | 10 | 6421 | 1401 | 2179 | 18597 | 74088 | 1142 | 642 | |
| Na | 11 | 6734 | 1146 | 2113 | 15725 | 586116 | 988 | 1510 | |
| Mg | 12 | 6204 | 1869 | 1930 | 11669 | 433911 | 805 | 1061 | |
| Al | 13 | 5745 | 2462 | 1788 | 9041 | 275293 | 675 | 615 | |
| Si | 14 | 5351 | 2953 | 1686 | 7281 | 187721 | 580 | 818 | |
| Р | 15 | 5022 | 3320 | 1630 | 6118 | 130332 | 516 | 2056 | |
| S | 16 | 4711 | 3726 | 1529 | 5049 | 98381 | 447 | 2203 | |
| Cl | 17 | 4441 | 4078 | 1443 | 4242 | 76636 | 391 | 2336 | |
| Ar | 18 | 4215 | 4339 | 1400 | 3703 | 60187 | 352 | 2975 | |

Since (1) is analytically integrable we obtain simply

$$\left\langle \frac{1}{r_{12}} \right\rangle_{\text{exch}} = N_8^5 \pi a_0 \sum_{l=1}^3 \frac{\gamma_l}{\sqrt{\lambda_l}} \,. \tag{5}$$

The exchange energies determined by the expression (5), and the corresponding HF SCF values (Kim & Gordon, 1974) are compared for a few atoms in Table 2.

 Table 2. Exchange energy values (a.u.) determined by

 the sum-rule procedure (SR) and calculated on the basis
 of the HF SCF atomic model

| | Z | SR | HF SCF |
|----|----|--------|--------|
| He | 2 | 1.0256 | 1.0387 |
| В | 5 | 3.7961 | 3.822 |
| N | 7 | 6.6743 | 6.742 |
| Ne | 10 | 12.125 | 12.60 |
| Ar | 18 | 30.21 | 30.09 |

Ions with atomic numbers 3 to 30

It has been found that better fits could be attained by separating the ions in two groups for which power n in (1) was fixed differently. For the ions with $2 < Z \le 20$, n=4, for those with $21 \le Z \le 29$, n=1 proved to be optimal.

Numerical values of the incoherent intensities of ions with $2 < Z \le 20$ have been taken from the *International Tables for X-ray Crystallography* (1962) and those for Z > 20 from the paper of Freeman & Watson (1961).

Tables 3 and 4 show the best parameters found for the two groups of ions and the percentage errors of the fits. The exchange energies as calculated from the analytic expression, contain a certain error especially in the case of heavy ions, probably because the fit is poorer in the range above $s > 1 \cdot 1$ Å⁻¹.

Table 3. Parameters $(\times 10^4)$ of the analytic fits of S(s) for ions with atomic numbers 3 to 20

. .

| $n = 4, \ s_{\max} = 1.1.$ | | | | | | | | | |
|----------------------------|----|------------|------|------|------|---------------|-------------|------|--|
| | Ν | Y 1 | Y2 | үз | λι | λ_2 | λ_3 | 3 | |
| Li+ | 2 | 55 | 7881 | 2066 | 2522 | 1469 2 | 23912 | 216 | |
| N- | 8 | 2396 | 3790 | 3801 | 2161 | 141843 | 37312 | 3231 | |
| 0- | 9 | 2275 | 1777 | 5948 | 1714 | 189366 | 36593 | 97 | |
| O+ | 7 | 2815 | 0 | 7201 | 1624 | 71049 | 34317 | 1921 | |
| F- | 10 | 2340 | 1440 | 6226 | 1643 | 168860 | 29172 | 2587 | |
| Na+ | 10 | 2144 | 1189 | 6667 | 877 | 46793 | 13597 | 303 | |
| Al+ | 12 | 2016 | 1868 | 6115 | 771 | 244040 | 9016 | 752 | |
| Al ³⁺ | 10 | 2176 | 887 | 6939 | 619 | 24640 | 8341 | 163 | |
| Si ³⁺ | 11 | 2027 | 1148 | 6827 | 526 | 125242 | 7047 | 575 | |
| Si ⁴ + | 10 | 2360 | 428 | 7217 | 608 | 23663 | 7248 | 460 | |
| Cl- | 18 | 4794 | 1956 | 3249 | 2152 | 186392 | 37255 | 2404 | |
| K+ | 18 | 4831 | 1515 | 3654 | 1629 | 93534 | 25374 | 812 | |
| Ca+ | 19 | 4677 | 1025 | 4298 | 1472 | 308314 | 25169 | 623 | |
| Ca ²⁺ | 18 | 4866 | 1236 | 3889 | 1444 | 78304 | 22023 | 295 | |
| Zn ²⁺ | 28 | 3291 | 1728 | 4981 | 540 | 43480 | 6047 | 123 | |
| | | | | | | | | | |

Table 4. Parameters $(\times 10^4)$ of the analytic fits of S(s) for ions with atomic numbers 21 to 29

 $n = 1, s_{\max} = 1 \cdot 1.$

| | Ν | 7 1 | Y2 | 7 3 | λ1 | λ_2 | λ_3 | 3 |
|-------------------|----|------------|------|------------|------|---------------|-------------|------|
| Sc+ | 20 | 3299 | 862 | 5842 | 5255 | 752363 | 90092 | 1664 |
| Sc ²⁺ | 19 | 3653 | 251 | 6097 | 5632 | 646734 | 98386 | 1625 |
| Sc ³⁺ | 18 | 3836 | 595 | 5578 | 5467 | 95597 | 95603 | 1864 |
| Ti+ | 21 | 2778 | 1346 | 5880 | 3892 | 500096 | 65939 | 1473 |
| Ti ²⁺ | 20 | 3051 | 689 | 6266 | 4105 | 499864 | 71796 | 1296 |
| Ti ³⁺ | 19 | 3458 | 754 | 5787 | 4593 | 169794 | 75246 | 1093 |
| V + | 22 | 2265 | 1809 | 5931 | 2656 | 365028 | 48736 | 954 |
| V ²⁺ | 21 | 2536 | 1727 | 5740 | 3037 | 229996 | 49774 | 1085 |
| V ³⁺ | 20 | 2962 | 1239 | 5801 | 3592 | 172320 | 56594 | 862 |
| Cr+ | 23 | 1898 | 1940 | 6166 | 1875 | 322099 | 39741 | 839 |
| Cr ²⁺ | 22 | 2112 | 2144 | 5748 | 2184 | 194745 | 38349 | 699 |
| Cr ³⁺ | 21 | 2429 | 2058 | 5516 | 2553 | 141207 | 40721 | 626 |
| Mn+ | 24 | 1433 | 2259 | 6312 | 682 | 271435 | 31573 | 899 |
| Mn ²⁺ | 23 | 1651 | 2340 | 6012 | 1245 | 175694 | 30415 | 592 |
| Mn ³⁺ | 22 | 1850 | 2724 | 5428 | 1468 | 119970 | 30747 | 541 |
| Fe+ | 25 | 1160 | 2360 | 6484 | 0 | 246356 | 37247 | 683 |
| Fe ²⁺ | 24 | 1242 | 2663 | 6096 | 264 | 152252 | 24820 | 467 |
| Fe ^{3 +} | 23 | 1271 | 3390 | 5340 | 382 | 98740 | 20773 | 478 |
| Co+ | 26 | 1080 | 2341 | 6583 | 0 | 233854 | 25106 | 776 |
| Co ²⁺ | 25 | 1116 | 2588 | 6299 | 55 | 147509 | 23001 | 509 |
| Co ³⁺ | 24 | 1273 | 3040 | 5688 | 495 | 99993 | 21071 | 262 |
| Ni+ | 27 | 989 | 2366 | 6641 | 0 | 216858 | 22660 | 875 |
| Ni ²⁺ | 26 | 2794 | 2624 | 4558 | 4781 | 124559 | 29215 | 2046 |
| Ni ³⁺ | 25 | 1893 | 2960 | 5147 | 2587 | 9 3298 | 22446 | 1130 |
| Cu+ | 28 | 862 | 2481 | 6660 | 23 | 193041 | 19928 | 1080 |
| Cu ²⁺ | 27 | 3053 | 619 | 6342 | 4678 | 668242 | 40843 | 2830 |
| Cu ³⁺ | 26 | 3001 | 717 | 6289 | 4284 | 207931 | 37378 | 1719 |
| | | | | | | | | |

Discussion

We have given an expression of good analytic behaviour for the approximation of incoherent X-ray and electron intensities of light atoms and ions. For the first 18 atoms this approximation is more accurate than the previous ones, and fits the numerical values uniformly over a wide range of the scattering variable.

For the ions, accurate approximations of the incoherent X-ray intensities have not been published up to the present. The present work is primarily intended for use in X-ray and electron diffraction studies on ionic solutions.

In both cases the mean errors are of the same order of magnitude, or even smaller than those of the fits of the corresponding coherent intensities.

Note: Two papers have been published meanwhile on the analytic approximation for the incoherent scattering function of atoms (Smith, Thakkar & Chapman, 1975; Balyuzi, 1975). The merit of the first one is to give an analytic formula of correct asymptotic behaviour, approximating the incoherent scattering functions of atoms over a wide s range. It is easy to show, that our analytic expression (1) has also a correct asymptotic behaviour at both high and low values of s. It can be verified, by comparing the corresponding percentage errors (ɛ) pertaining to atoms common in both papers, that the present approximation is a better fit. The approximating formula, however, referred to in the second paper is valid only up to s = 1.5, and its asymptotic behaviour is not correct. Standard deviations given by the author have to be treated with criticism because they relate to the function $Z-I_{inc}(s)$ instead of the incoherent intensities themselves. Having converted them to the suitable form, we could establish that the percentage errors (ε) in the range $0 \le s \le 1.5$ were similar to ours.

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